

Super-Kamiokande (SK) has been and continues to be one of the most successful experiments in neutrino physics. However, SK cannot distinguish neutrinos from anti-neutrinos, a crucial capability for several very important SK measurements. Being able to do so would improve dramatically SK's sensitivity to Supernova Relic Neutrinos since its currently overwhelming spallation background could be largely suppressed, and it would also enable a precise determination of solar elements of the leptonic mixing matrix from reactor neutrinos.

For that purpose SK is carrying out EGADS, a GADZOOKS! R&D project. The idea is (Beacom and Vagins, 2004) is to dissolve in the water a compound of Gadolinium (Gd), the element that has the largest cross section for neutron capture, process that is followed by the emission of few gammas adding up 8 MeV energy. The Gd would allow SK to identify anti-neutrinos interacting via inverse beta decay, by coincident signals from the charged anti-lepton and the delayed 8 MeV  $\gamma$  cascade.

This technique could have a serious drawback if the Gd compound brings with it a sizable contamination of natural radioactivity. We are thoroughly studying this potential problem by analyzing the different background-producing final states induced by the spontaneous fission of the  $^{238}\text{U}$  and by the  $\alpha$  particles from the  $\alpha$ -decays in the decay chains. We are estimating the amount of radioactive contamination that SK with Gd can afford, and are investigating reconstruction algorithms that could maximize the discrimination between anti-neutrino signal and radioactivity-induced background.

## Introduction

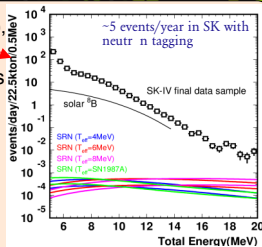
- SK is a 50 kton water Cherenkov detector in the Kamioka Mine, Japan
- SK began its search for proton decay and neutrino physics studies in 1996
- Important is the CCQE scattering of the incoming  $\nu$  with nucleons in water



- Nowadays, SK cannot distinguish (efficiently)  $\bar{\nu}$  from  $\nu$  because it is not able to detect either neutrons nor protons (the latter at low energy)
- This distinction would be of most importance for deeper study of  $\nu$  properties  $\rightarrow$  try neutron tagging

## New physics reach

- Diffuse Supernova Neutrino Background (DSNB), currently limited by background  $\nu_e + p \rightarrow e^+ + n$
- Improve sensitivity to solar sector oscillation parameters  $\rightarrow$  neutron tagging would increase substantially the rate compared to the KamLAND reactor
- proton decay: no neutron should be in the final state
- Sensitivity to "wrong sign"  $\nu_e$  production from Sun
- Much more detailed measurement of nearby Supernovae explosion
- Gain information on the hadronic final states of  $\nu$ s interactions
- Maybe distinction between  $\bar{\nu}$ s and  $\nu$ s from T2K
- Identification of  $\bar{\nu}$  over  $\nu$  in the atmospheric sector ( $\sim$ GeV) is being studied

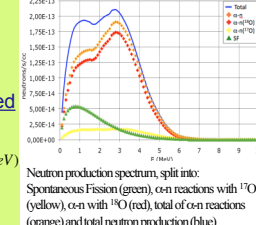


## Backgrounds of GADZOOKS!

- An important source of background comes from the radioactive contamination of the materials; of clear relevance is that in the Gd...
- We study thoroughly these backgrounds:
  - their intensity given the measured radioactivities
  - their impact on the new physics measurements
  - their impact on the traditional SK measurements
- the complete campaign of radioactivity for GADZOOKS! is carried out at the LSC (Canfranc Underground Laboratory) using HPGe detectors

### neutrons from $(\alpha, n)$

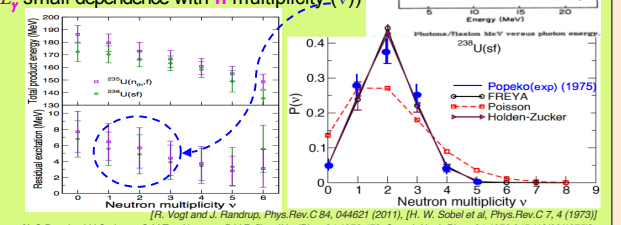
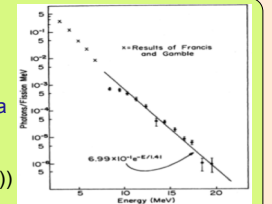
- $(\alpha, n)$  reactions with  $^{18}\text{O}$ 
  - $\alpha$  particles are emitted with  $\sim 6$  MeV
  - $^{18}\text{O} + \alpha (\sim 6\text{MeV}) \rightarrow ^{22}\text{Ne}^* \rightarrow ^{20}\text{Ne} + 2n (\sim 2.5\text{MeV})$
  - $\rightarrow$  neutrons come in pairs and no  $\gamma$ s are emitted
- Similarly occurs with  $^{17}\text{O}$  case, but
  - $^{17}\text{O} + \alpha (\sim 6\text{MeV}) \rightarrow ^{21}\text{Ne}^* \rightarrow ^{20}\text{Ne} + n + \gamma (0.4\text{MeV}) + \beta (1.5\text{MeV})$
  - $\rightarrow$  neutrons come in pairs
  - $\rightarrow$   $\gamma$ s and  $\beta$ s are too weak to be detectable



$(\alpha, n)$  are distinguishable from inverse beta... But produce many neutrons

### $[1 + 1n]$ events from Spontaneous Fission

- $^{238}\text{U}$  the greatest contribution to SF
  - emits neutrons and photons
  - fraction of SF in  $^{238}\text{U}$  is  $5 \cdot 10^{-7}$
  - probability of a  $\gamma$  ray from  $^{238}\text{U}$  SF with a given energy is  $P(E) = 0.7e^{-E/1.4}$
  - probability of 1n per fission is 28% ( $E_\gamma$  small dependence with n multiplicity ( $\nu$ ))



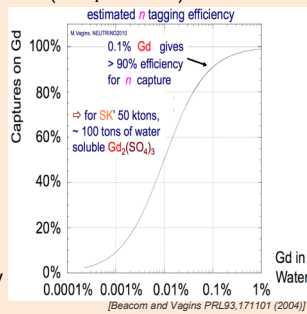
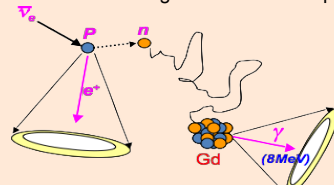
- $\rightarrow 60 [\gamma + n]$  events/year
- AJ4400 reduces the amount of  $^{238}\text{U}$  by a factor 100
- $\rightarrow 0.6 [\gamma + n]$  events/year after AJ4400 processing
- current  $\text{Gd}_2(\text{SO}_4)_3$  batch at EGADS  $\rightarrow < 2 [\gamma + n]$  events/year with AJ4400

### s and $\beta$ s

- Other elements of radioactive chains decay by emitting  $\beta$ s
  - Some elements emit  $\beta$ s with sufficiently high energy accompanied by  $\gamma$ s
- $^{232}\text{Th} \rightarrow \dots \rightarrow ^{228}\text{Ra} \rightarrow \dots \rightarrow ^{208}\text{Tl} \rightarrow E_{\beta, \text{max}} = 5.0 \text{ MeV} + \begin{cases} 583 \text{ keV } 84\% \\ 861 \text{ keV } 12\% \\ 2615 \text{ keV } 99\% \end{cases}$
- $^{238}\text{U} \rightarrow \dots \rightarrow ^{226}\text{Ra} \rightarrow ^{222}\text{Rn} \rightarrow \dots \rightarrow ^{214}\text{Bi} \rightarrow E_{\beta, \text{max}} = 3.3 \text{ MeV} + \begin{cases} 609 \text{ keV } 46\% \\ 1120 \text{ keV } 15\% \\ 1765 \text{ keV } 16\% \end{cases}$
- To achieve 4.0 MeV solar neutrino analysis in 0.1%Gd+H<sub>2</sub>O SK
  - $< 0.35 \text{ mBq/kg } [^{238}\text{U}]$
  - $< 0.35 \text{ mBq/kg } [^{232}\text{Th}]$
- we also need to remove the Ra from Gd compound
- Selective resins, such as DOWEX RSC, are one available solution

## GADZOOKS! project

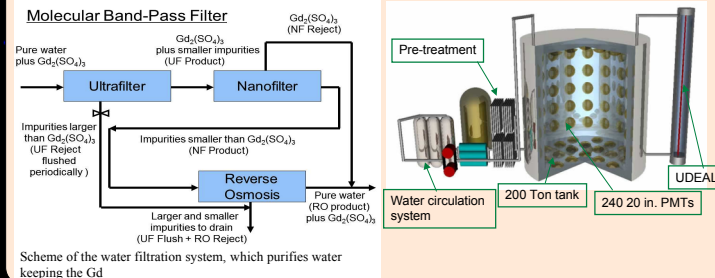
- neutron tagging in Super-Kamiokande
- Gadolinium has the greatest n capture cross section of all stable nuclei
- When Gd captures a thermalized n, it emits a  $\gamma$  cascade of 8 MeV
- SK will be able to tag the inverse beta process ( $\nu_e + p \rightarrow e^+ + n$ ):



- A water soluble Gd compound added to WC detectors would greatly improve their performance via delayed coincidence
- $\text{Gd}_2(\text{SO}_4)_3$ : no corrosion, excellent solubility and small light attenuation

## EGADS

- Selective filtration system (brand new technology developed at UCI)
- Gd compound "cleaning" and dissolving  $\rightarrow$  pre-treatment system circulation thorough resins (AJ4400 to remove U, Th)
- Water transparency monitoring  $\rightarrow$  UDEAL (Underground Device Evaluating Attenuation Length)
- Check Gd volume uniformity  $\rightarrow$  AAS (Atomic Absorption Spectrometer)
- Gd removal



## Summary / Conclusions

- Neutron tagging would improve dramatically the sensitivity of SK
- Gd very good candidate to do the job
- $\text{Gd}_2(\text{SO}_4)_3$  best compound in terms of solubility and corrosion
- Tests done with  $\text{Gd}_2(\text{SO}_4)_3$  show
  - no Gd rejection due to the selective filtration system: success!
  - very good water transparency
  - solution uniformity very good
- Studied backgrounds from radioactive contamination
  - $(\alpha, n)$  produce many n but no high energy  $\gamma$ s
  - background from  $^{238}\text{U}$  Spontaneous Fission implies that we must reduce U contamination in Gd salt. AJ4400 removes U with high enough efficiency
  - $\beta$ s and  $\gamma$ s from  $^{208}\text{Tl}$  and  $^{214}\text{Bi}$  implies we must reduce Ra levels